

Professor Jasper Halekas
Van Allen 70
MWF 12:30-1:20 Lecture

## Announcements

- First labs start today!
- First homework assignment due in class (or before) on Friday


## Last Time Recap

- Events that are simultaneous in one inertial frame are not necessarily simultaneous in another inertial frame
- The time duration measured between two events is different in different frames
- $\Delta t=\gamma \Delta \tau$
" $\Delta \tau=\Delta t_{0}$ is the "proper time" measured in a frame where the two events take place at the same location


## Back to Michelson-Morley



Length contraction
$\xrightarrow[\substack{L+\\ H_{L}^{L}}]{\stackrel{u}{L}}$

$$
\begin{aligned}
c \Delta t_{1} & =L+u \Delta t_{1} \\
c \Delta t_{2} & =L-u \Delta t_{2} \\
\Delta t & =\Delta t_{1}+\Delta t_{2} \\
& =\frac{L}{c-u}+\frac{L}{c+u} \\
& =2 L / c \cdot \frac{1}{1-u^{2} / c^{2}} \\
& =2 L / c-\gamma^{2}
\end{aligned}
$$

But $\Delta t=\gamma \Delta t^{\prime}=\gamma \Delta t_{0}$

$$
w / \Delta t_{0}=2 L_{0} / c
$$

= time interval in Earth frame

$$
\begin{aligned}
& \Rightarrow 2 L / c \cdot \gamma^{2}=2 L_{0} / c \cdot \gamma \\
& \Rightarrow \gamma L=L_{0} \\
& \Rightarrow L=L_{0} / \gamma \\
& w / L_{0}=\text { "proper length" }
\end{aligned}
$$

## Proper Length

- "Proper Length" = L。
- The length of an object measured in a frame where it is at rest
- The length $L$ measured in any other frame moving with a velocity with respect to this frame that has a component along the length will be shorter
- L = $L_{0} / \gamma$


## Length Contraction


$\mathrm{V}=0$


## Concept Check: Proper Time/Length

- A star (assumed to be at rest relative to the Earth) is 100 light-years from Earth. (A light-year is the distance light travels in one year.) An astronaut sets out from Earth on a journey to the star at a constant speed of o.98c. (Note: At $u=0.98 c, \gamma=5$ )
- How long does a light signal take to reach the star, and how long does it take the astronaut to reach the star, both according to Earth?
A. 100 years, 98 years
B. 98 years, 100 years
C. 100 years, 102 years
D. 100 years, 100 years


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| :--- |

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- According to the astronaut, what is the distance to the star, and how long does it take to get there?
A. 100 light years, 100 years
B. 20 light years, 20 years
C. 20 light years, 20.4 years
D. 100 light years, 20.4 years
E. 20 light years, 102 years


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## Concept Check: Proper Time/Length

- Light takes 100 years to travel to the star, but according to the astronaut he makes it there in 20.4 years. Does that mean he travels faster than light?
A. Yes
B. No
C. Maybe?
D. WTF?


## Concept Check: Proper Time/Length

- Light takes 100 years to travel to the star, but according to the astronaut he makes it there in 20.4 years. Does that mean he travels faster than light?
A. Yes
B. No
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## Moral of the story

The proper time and the proper length do not have to be in the same frame!

- Astronaut says he travels distance $\mathrm{L}=\mathrm{L}_{0} / \gamma$ in proper time $\mathrm{t}_{\mathrm{o}}$
- From Earth, we see astronaut travel proper length $L_{0}$ in time $t=\gamma t_{0}$
- A muon's trip through the atmosphere is a similar story


## The Lorentz Transformation: Preamble



A stick with length $L_{o}$ is at rest in $S^{\prime}$. Its endpoints are the events $(x, t)=(0,0)$ and $\left(x^{\prime}, 0\right)$ in $\mathrm{S}^{\prime}$. $\mathrm{S}^{\prime}$ is moving to the right with respect to frame $S$ at a velocity $u$.

Event 1 - left of stick passes origin of S. Its coordinates are ( 0,0 ) in S and ( 0,0 ) in $\mathrm{S}^{\prime}$.

## Concept Check: Lorentz Transformation

An observer at rest in frame $S$ sees a stick flying past him with velocity u:


As viewed from S , the stick's length is $L_{o} / \gamma$. Time t passes. According to $S$, where is the right end of the stick? (Assume the left end of the stick was at the origin of $S$ at time $t=0$.)
A) $x=\gamma u t$
B) $x=u t+x^{\prime} / \gamma$
C) $x=-u t+x^{\prime} / \gamma$
D) $x=u t-x^{\prime} / \gamma$
E) Something else...

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As viewed from S , the stick's length is $L_{o} / \gamma$. Time t passes. According to $S$, where is the right end of the stick? (Assume the left end of the stick was at the origin of $S$ at time $t=0$.)
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D) $x=u t-x^{\prime} / \gamma \quad$ E) Sometning else...

Larent $z$ Transformation

$$
\Rightarrow \begin{aligned}
& x=u t+x^{\prime} / r \\
& \Rightarrow x^{\prime}=\gamma\left(x^{\prime}-u t\right) \\
& \varphi^{\prime}=\gamma \\
& z^{\prime}=z \\
& t^{\prime}=\gamma\left(t-u / c^{2} x\right)
\end{aligned}
$$

- Align origins

$$
\begin{aligned}
& {[x, y, z, t] } \\
= & {[x, y, z, t] } \\
= & {[0,0,0}
\end{aligned}
$$

when left end of stick (a) origin

- At time t

$$
\begin{aligned}
& x_{L}=u t \\
& x_{L}=\gamma(u t-u t)=0 \\
& x_{R}=u t+L_{0} / \gamma \\
& x_{R}^{\prime}=\gamma\left(u t+L_{0} / \gamma-u t\right)=L_{0} \\
& \Delta x_{0}=x_{R}-x_{L}=L_{0} / \gamma \\
& \Delta x^{\prime}=L_{0}
\end{aligned}
$$

$$
\begin{aligned}
-t_{L}^{\prime} & =\gamma\left(t-u / c^{2} x_{L}\right)=r\left(t-u^{2} / L^{2} t\right) \\
t_{R}^{\prime} & =r\left(t-u / c^{2} x_{R}\right)=r\left(t-u^{2} / c^{2} t-u / c^{L} / / r\right) \\
\Delta t^{\prime} & =-u^{2} / L_{0} L_{0}
\end{aligned}
$$

- clocks out of syne according to moving observer


## Transformations

If $S^{\prime}$ is moving with speed $v$ in the positive $x$ direction relative to $S$, then the coordinates of the same event in the two frames are related by:

## Galilean transformation (classical)

$$
\begin{aligned}
& x^{\prime}=x-u t \\
& y^{\prime}=y \\
& z^{\prime}=z \\
& t^{\prime}=t
\end{aligned}
$$

Lorentz transformation (relativistic)

$$
\begin{aligned}
x^{\prime} & =\gamma(x-u t) \\
y^{\prime} & =y \\
z^{\prime} & =z \\
t^{\prime} & =\gamma\left(t-\frac{u}{c^{2}} x\right)
\end{aligned}
$$

Note: This assumes ( $0,0,0,0$ ) is the same event in both frames.

## Interval Transformations

If $S^{\prime}$ is moving with speed $v$ in the positive $x$ direction relative to $S$, then the coordinates of the same event in the two frames are related by:

Galilean transformation (classical)
$\Delta x^{\prime}=\Delta x-u \Delta t$
$\Delta y^{\prime}=\Delta y$
$\Delta z^{\prime}=\Delta z$
$\Delta t^{\prime}=\Delta t$

Lorentz transformation
(relativistic)

$$
\begin{aligned}
& \Delta x^{\prime}=\gamma(\Delta x-u \Delta t) \\
& \Delta y^{\prime}=\Delta y \\
& \Delta z^{\prime}=\Delta z \\
& \Delta t^{\prime}=\gamma\left(\Delta t-\frac{u}{c^{2}} \Delta x\right)
\end{aligned}
$$

## Relativistic Doppler Shift



Relativistic Doppler Shift


Frequency $f^{\prime} Q$ origin

|  | $s^{\prime}$ | $s$ |
| :--- | :--- | :--- |
| Pulse 1 | $x_{1}^{\prime}=0$ | $x_{1}=0$ |
|  | $t_{1}^{\prime}=0$ | $t_{1}=0$ |
| Pulse 2 | $x_{2}^{\prime}=0$ | $x_{2}=\gamma u t_{2}^{\prime}$ |
|  | $t_{2}^{\prime}=1 / f^{\prime}$ | $t_{2}=\gamma t_{2}^{\prime}$ |

-Pulse 2 takes $\Delta t=\frac{x_{2}-x_{1}}{c}$ to reach origin in $S$

- Pulse 2 received at

$$
\begin{aligned}
1 / f & =t_{2}-t_{1}+\Delta t \\
& =\gamma t_{2}^{\prime}+\gamma u t_{2}^{\prime} / c \\
& =\gamma(1+w / c)+2 \\
& =\gamma(1+w / c)-\gamma / f^{\prime}
\end{aligned}
$$

or $f=\frac{1}{\gamma} \frac{1}{1+w / c} f^{\prime}$

$$
\begin{aligned}
f & =\frac{\sqrt{1-u^{2} / c^{2}}}{1+u / c} f^{\prime} \\
& =\frac{\sqrt{(1+u / c)(1-u / c)}}{1+u / c} f^{\prime} \\
f & =\sqrt{\frac{1-u / c}{1+u / c}} f^{\prime}
\end{aligned}
$$

- (an write as

$$
\begin{gathered}
f=\sqrt{\frac{1-\beta}{1+\beta}} f^{\prime} \\
w / \beta=u / c
\end{gathered}
$$

## Relativistic Doppler Shift



